

Here are some comments from me, before others' work:

3. This was mostly good. We had anticipated starting with $GM = 4 \times 10^{14}$ for the earth and using 15% of that. I checked current reported values and they seem to indicate that Mars is actually closer to 10.8% of the Earth, not 15%. Please do not use this for planning your mission to Mars.

4. - 5. were mostly good also. To be clear, I checked data here also. The mean radius of the Earth is about 6.371×10^6 m. Rounding that to 6.4 and adding 300 km gives the book number of 6.7×10^6 m for distance from the centre of the Earth for the rocket.

8. This was the trickiest one. Most importantly, this is a mathematics class so I need a mathematical justification for your answer. Quoting physics formulae or saying "make this negative" or "use the other value" does not suffice. Both rocket-firings, the beginning and the second should actually increase the speed. So, $v_1 < 7785$ and $v_2 < 7560$. (It is true that we could have computed these circular speeds from the given information. I did, and they checked.) Please read the exemplar and see how the steps follow. A summary is: use apogee and perigee to solve for eccentricity and gamma. From gamma, find K. Once you know K, you can easily find v from r, since at apogee and perigee r and v are perpendicular.

7. There needs to be comprehensible work for computing your line integrals. I like to see pullbacks, but that isn't needed. I do think actual integration is needed, because it isn't a constant form.

16. The best way is demonstrated in the exemplar, there are others. Notice the question says "where along this segment". A t-value doesn't answer that, we need ranges of points on the segment itself.

3.4

$$(3) \quad r = 3300 \text{ km}, \quad M_M = 0.15 m_e$$

$$GM_e = 4 \cdot 10^{14} \text{ m}^3/\text{s}^2 \rightarrow GM_M = 6 \cdot 10^{13} \text{ m}^3/\text{s}^2$$

Using work done from pages 70-71,

$$|\epsilon| > 1$$

$$e^2 = \sin^2 \phi \left(1 - \frac{rv^2}{GM} \right) + \cos^2 \phi$$

$$v_s > \sqrt{\frac{2GM}{r}} \rightarrow \text{independent of } \phi, \text{ "escape speed"}$$

↳ doesn't sound as smart

$$v_s > \sqrt{\frac{1.2 \cdot 10^{14}}{3.3 \cdot 10^6}}$$

$$v_s > 6030.23 \text{ m/s}$$

Consider a rocket fired 300 km above earth's surface
 6.7×10^6 m from the center.

4.) What velocity must it attained if it is to remain in a circular orbit at this height? What is the period of this orbit?

$$V_0 = \sqrt{\frac{GM}{R}} \quad \text{where } R \text{ is the Radius of orbit}$$

$$\text{Let } GM = 4 \times 10^{14} \text{ m}^3/\text{s}^2 \quad R = 6.7 \times 10^6 + 3 \times 10^5 = 7,000,000$$

$$V_0 = \sqrt{\frac{4 \times 10^{14}}{6.7 \times 10^6}} = 7726.67 \text{ m/s}$$

$$\text{Period} = \frac{2\pi(6.7 \times 10^6)}{7726.67} = 5448.31 \text{ seconds} = 90 \text{ minutes and } 48 \text{ seconds}$$

#5. Speed is 9000 m/s and $\phi = \pi/2$. What are the values of the apogee and the perigee?

Since we know the angle b/w \vec{r} & \vec{v} is denoted by ϕ , then

$$K = r v \sin \phi$$

We know $r = 6.7 \text{ E}6$

$$\text{thus, } K = (6.7 \text{ E}6)(9000) \sin(\pi/2) = 6.03 \text{ E}10$$

$$\text{By equation (3.41)} \quad \gamma = \frac{K^2}{GM} = \frac{(6.03 \text{ E}10)^2}{(4 \text{ E}14)} = 9.09 \text{ E}6$$

Finding ϵ by using eqn (3.54)

$$\epsilon = \frac{rv^2}{GM} - 1 = \frac{(6.7 \text{ E}6)(9000)^2}{(4 \text{ E}14)} - 1 = 0.35675$$

thus

$$\text{apogee} = \frac{\gamma}{1 - |\epsilon|} = \frac{9.09 \text{ E}6}{1 - 0.35675} = 1.413 \text{ E}7$$

$$\text{perigee} = \frac{\gamma}{1 + |\epsilon|} = \frac{9.09 \text{ E}6}{1 + 0.35675} = 6.6998 \text{ E}6$$

$$\text{period} = \frac{2\pi}{\sqrt{GM}} a^{3/2} = \frac{2\pi (1.041 \text{ E}7)^{3/2}}{\sqrt{(4 \text{ E}14)}} = 10560.3$$

$$a = \frac{\gamma}{1 - \epsilon^2} = \frac{9.09 \text{ E}6}{1 - (0.35675)^2} = 1.041 \text{ E}7$$

$$8) r_i = 6.6 \times 10^6 \quad v_i = 7785$$

$$= 7.0 \times 10^6 \quad v_i = ?$$

$$\text{apogee} = 7 \times 10^6 \quad v_a = ?$$

$$r_f = 7.0 \times 10^6 \quad v_f = 7560$$

$$\text{perigee} = 6.6 \times 10^6$$

$$6.6 \times 10^6 (1 + \epsilon) = \gamma$$

$$7 \times 10^6 (1 - \epsilon) = \gamma$$

$$6.6 \times 10^6 (1 + \epsilon) = 7 \times 10^6 (1 - \epsilon)$$

$$1 + \epsilon = \frac{35}{33} (1 - \epsilon)$$

$$1 + \epsilon = \frac{35}{33} - \left(\frac{35}{33}\right) \epsilon$$

$$1 + \left(\frac{68}{33}\right) \epsilon = \frac{35}{33}$$

$$\left(\frac{68}{33}\right) \epsilon = \frac{2}{33}$$

$$\epsilon = 0.0294$$

$$K = r v \sin \phi$$

$$5.21 \times 10^{10} = 6.6 \times 10^6 v (\sin \pi/2)$$

$$v_i = 7893.94 \text{ m/s}$$

$$K = r v \sin \phi$$

$$5.21 \times 10^{10} = 7 \times 10^6 v (\sin \pi/2)$$

$$v_a = 7442.86$$

$$\gamma = 6.6 \times 10^6 (1.0294)$$

$$= 6794040$$

$$\gamma = \frac{K^2}{GM}$$

$$6794040 = \frac{K^2}{4 \times 10^{14}}$$

$$K = 5.21 \times 10^{10}$$

$$v_i = 7893.94 \text{ m/s}$$

$$v_a = 7442.86 \text{ m/s}$$

- Steps
- 1) find $\vec{r}(t)$
 - 2) compute r^* (differential form)
↳ replace x, y, z from $\vec{r}(t)$
 - 3) solve

4.3 #7) $y dx + z dy + x dz$

pullback

a) from $(0, 2, -1)$ to $(3, 0, 1)$

$$\vec{r}(t) = \vec{a} + t(\vec{b} - \vec{a})$$

$$= (0, 2, -1) + t(3, -2, 2)$$

$$= \left(\underbrace{3t}_x, \underbrace{2-2t}_y, \underbrace{-1+2t}_z \right)$$

$\vec{r}^{-1}(b)$

① $\int r^*(y dx + z dy + x dz)$

$\vec{r}^{-1}(a)$

$r^*(y) dx = (2-2t) \cdot 3 dt$ $r^*(z) dy = -1+2t(-2 dt)$ } combine these in ①

$r^*(x) dz = 3t \cdot 2 dt$

$$\int_0^1 (6 - 6t + 2 - 4t + 6t) dt$$

$$= \int_0^1 (8 - 4t) dt$$

$$= 8t - 2t^2 \Big|_0^1 = 8 - 2 = \boxed{6}$$

b) from $(3, 0, 1)$ to $(0, 2, -1)$

step 1

$\vec{r}(t) = (3, 0, 1) + t(-3, 2, -2)$

$$= (3-3t, 2t, 1-2t)$$

$x = 3-3t$
 $y = 2t$
 $z = 1-2t$

step 2

$\int_0^1 r^*(y dx + z dy + x dz)$

$r^*(y) dx = 2t \cdot -3 dt$

$r^*(z) dy = (1-2t) \cdot 2 dt$

$r^*(x) dz = (3-3t) \cdot -2 dt$

step 3

$$\int_0^1 (-6t + 2 - 4t - 6 + 6t) dt$$

$$= \int_0^1 (-4 - 4t) dt$$

$$= -4t - 2t^2 \Big|_0^1 = -4 - 2 = \boxed{-6}$$

c) from $(3, 1, 2)$ to $(-1, 1, 1)$

step 1

$\vec{r}(t) = (3, 1, 2) + t(-4, 0, -1)$

$$= (3-4t, 1, 2-t)$$

$x = 3-4t$
 $y = 1$
 $z = 2-t$

step 2

$\int_0^1 r^*(y dx + z dy + x dz)$

$r^*(y) dx = 1 \cdot -4 dt$ $r^*(z) dy = (2-t) \cdot 0 dt$ $r^*(x) dz = (3-4t) \cdot -1 dt$

step 3

$$\int_0^1 (-4 - 3 + 4t) dt \rightarrow \int_0^1 (-7 + 4t) dt \rightarrow -7t + 2t^2 \Big|_0^1 = -7 + 2 = \boxed{-5}$$

7d) form: $ydxdz + zdy + xdz$
 from $(0, 2, -1)$ to $(1, 3, 2)$
 \vec{a} \vec{b}

step 1 $\vec{r}(t) = (0, 2, -1) + t(1, 1, 3)$ $x = t$
 $= (t, 2+t, -1+3t)$ $y = 2+t$
 $z = -1+3t$

step 2 $\int_C r^*(ydx + zdy + xdz)$

step 3 $r^*(y)dx = (2+t)dt$ $r^*(z)dy = (-1+3t)dt$ $r^*(x)dz = t \cdot 3dt$
 $= \int_0^1 (2+t - 1 + 3t + 3t) dt$
 $= \int_0^1 (7t + 1) dt \rightarrow 3.5t^2 + t \Big|_0^1 = 3.5 + 1 = \boxed{4.5}$

e) from $(1, 3, 2)$ to $(3, 0, 1)$
 \vec{a} \vec{b}

step 1 $\vec{r}(t) = (1, 3, 2) + t(2, -3, -1)$ $x = 1+2t$
 $= (1+2t, 3-3t, 2-t)$ $y = 3-3t$
 $z = 2-t$

step 2 $\int_C r^*(ydx + zdy + xdz)$

step 3 $r^*(y)dx = (3-3t)2dt$ $r^*(z)dy = (2-t) \cdot (-3)dt$ $r^*(x)dz = (1+2t) \cdot (-1)dt$
 $= \int_0^1 (6-6t - 3t + 1 - 2t) dt$
 $= \int_0^1 (-1 - 5t) dt \rightarrow -t - 2.5t^2 \Big|_0^1 = -1 - 2.5 = \boxed{-3.5}$

16. Fluid flow: $(y-y^2)dx + (xy-x)dy$

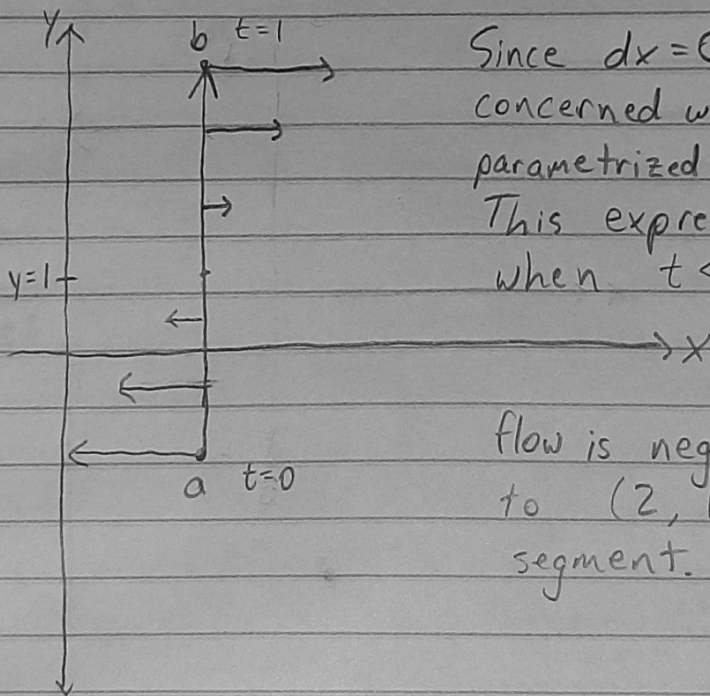
$$\vec{a} = (2, -2) \quad \vec{b} = (2, 4) \quad \text{Using } (x, y) = \vec{a} + t(\vec{b} - \vec{a})$$

$$\begin{aligned} x &= 2 & dx &= 0 \\ y &= -2 + 6t & dy &= 6dt \end{aligned}$$

$$(y-y^2)dx = [(-2+6t) - (4-12t+36t^2)](0) = 0$$

$$(xy-x)dy = (-4+12t-2)6dt$$

$$\int_0^1 (-36+72t)dt = -36t + 36t^2 \Big|_0^1 = \boxed{0}$$



Since $dx=0$, we are only concerned with the dy term, parametrized by $-36+72t$.

This expression is negative when $t < \frac{1}{2}$, so the

flow is negative from $(2, -2)$ to $(2, 1)$ along the line segment.